Combining Arrays



e-MERLIN+EVLA EVN+*e*-MERLIN Different EVLA or ALMA configurations



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Science & Technology

Filling the aperture - NGC 7027







Improving aperture (uv) coverage



- Less of a problem for new broad bands

- But still important for high-resolution/high frequency
- Vital for spectral line

August 2009

STFC Summer School A M S Richards

U Ori Masers MERLIN+EVN

- R_{OH mainline shell} ~3 MERLIN beams
- EVN resolves-out 10-90%
- Combine best of both!





Aligning data from two arrays

- Compare flux on overlapping baselines
- Check position
 - Ideally observed with same phase-reference source
 - Correct coords, UVFIX or self-cal:
 - Make image from high-res data, tapered to use short baselines in common with low-res data
 - Use to self-cal low-res data
 - Useful for bright sources with multiple peaks
- Frequency:
 - Lines: same spectral config., at fixed v if possible
 - Allow for spectral index if continuum at dif. $\boldsymbol{\nu}$
 - May combine different continuum configs
 - In map plane (dirty maps) or *uv* data in CASA
- Can add in single-dish data ('feathering')





Astronomical Image Processing System

- Originated by NRAO for VLA in 1978
 - Fortran, C
 - Limited built-in scripting/math operations
 - Recent python wrapper (Parseltongue)
 - Now most widely used package for cm-wave
 - VLA, MERLIN, most VLBI ... many more interferometers
 - Some support for single dish
 - Recognises other images e.g. HST, X-ray...
 - Very wide functionality from calibration to analysis
- Binary releases for Linux/Unix, Mac o/s
 - Local compilation possible but rarely needed



Flexible Image Transport System

- Standard astronomical data format:
 - UVFITS or IDI FITS for visibility data
 - Image files for 1, 2, 3+ D images
 - Unfortunately several dialects
 - AIPS uses FITS
 - CASA can read/export some FITS
- Structure of FITS file
 - Header
 - (Binary) data
 - Extension tables

- Fortunately there are tools • IMHEAD in AIPS or CASA



A M S Richards MWA 2010 p7 Printed: 22

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	The Astronomical Image and Table Format
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- Fortunately there are tools • IMHEAD in AIPS or CASA

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ALM/

AIPS overview

- Use recent version (Dec10 in these examples)
- Start: type >aips
- Starts an environment in your xterm
 - Additional TV, message and Tek (plot) windows
 - Try >aips tv=local if problems using TV
 - Occasionally (dis)connecting from internet messes up
 - Limited number of instances can be run at once
- Enter an AIPS Number >1 (make a note of it)
 - Convenient to use a different number per dataset
 - Multi-user systems may have allocated numbers



AIPS jargon

- Major operations are performed using Tasks
 FITLD loads data, CALIB performs calibration etc.
- Input parameters to Tasks are set by Verbs
 - >Task 'CALIB'; CALSOUR 'MKN273'; SOLINT 1
 - Words/names in 'inverted commas'; numbers bare
 - Not case sensitive, in general
 - Inside AIPS, 12-character limit on file/source names
- To set all defaults: >RESTORE 0
 - Beware: will give values typical for VLA data
 - You will have to set parameters suitable for your data
- To exit and kill all AIPS windows: >KLEENEX





Starting AIPS

EUROPEAN ARC

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[amsr@KALI INTERFERO]; aips tv=local START_AIPS: Will use or start first available Unix Socket based TV START_AIPS: User data area assignments: (Using global default file /home/amsr/aips/DA00/DADEVS.LIST for DADEVS.PL Disk 1 (1) is /home/amsr/aips/DATA/KALI_1	_)
START_AIPS: Starting TPMON daemons on KALI asynchronously Starting up 31DECO9 AIPS with normal priority Begin the one true AIPS number 1 (release of 31DECO9) at priority = 0 AIPS 1: You are not on a local TV device, welcome stranger AIPS 1: You are assigned TV device/server 2 AIPS 1: You are assigned oraphics device/server 2	
AIPS 1: Enter user ID Humber 289 AIPS 1: Copyright (C) 1995-2009 Associated Universities, Inc. AIPS 1: Copyright (C) 1995-2009 Associated Universities, Inc. AIPS 1: AIPS comes with ABSOLUTELY NO WARRANTY; AIPS 1: AIPS comes with ABSOLUTELY NO WARRANTY; AIPS 1: For details, type HELP GNUGPL AIPS 1: This is free software, and you are welcome to redistribute it AIPS 1: under certain conditions; type EXPLAIN GNUGPL for details. AIPS 1: Previous session command-line history recovered.	
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<pre>[amsr@KALI INTERFER0]\$ aips tv=local START_AIPS: Will use or start first available Unix Socket ba START_AIPS: Your initial AIPS printer is the START_AIPS: - system name , AIPS type START_AIPS: User data area assignments: (Using global default file /home/amsr/aips/DA00/DADEVS.LIS Disk 1 (1) is /home/amsr/aips/DATA/KALI_1 Tape assignments: Tape 1 is REMOTE Tape 2 is REMOTE Tape 2 is REMOTE START_AIPS: I am GUESSING you are at a workstation called ka START_AIPS: Starting TV servers on kali asynchronously START_AIPS: Starting TV servers on kali asynchronously START_AIPS: Starting TV servers on KALI asynchronously START_AIPS: Starting TPMON daemons on KALI asynchronously Starting up 31DECO9 AIPS with normal priority UNIXSERVERS: Start TV LOCK daemon TVSRV1 on kali Begin the one true AIPS number 1 (release of 31DECO9) at prior STARTPMON: [KALI] Starting TPMON1 with output SUPPRESSED UNIXSERVERS: Start XAS1 on kali, DISPLAY :0.0</pre>	<pre>AIPS_MSGSRV_1 MSGserver: Starting AIPS task logging, Unix (local) domain hostna> task #: Message </pre>	

AIPS

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ALMA ALMA Regional Centre || UK

Loading data into AIPS

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EUROPEAN ARC ALMA Regional Centre II UK

Loading data into AIPS



Where does AIPS put data?

AIPS_MSGSRV_1

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hostna> task #: Message \times 🗖 xterm > FITLD1: Task FIT KALI [amsr@KALI > FITLD1: Found Mk KALI. [amsr@KAL___]\$ ls /home/amsr/aips/DATA/KALI_] |KALI > FITLD1: Create M AND001001.01. CCD005046.00R: CCD00A067 224; FGD002001.05K: CCD00M01Y.00U: KALI > FITLD1 AND001001.023; CCD005017.02R, CCD00H06Z.ORQ; CCD00M01Z.00U: FGD002001.00U: > FITLD1: Check IKALI. CCD005048,00R; AND001001.02H: CCD00A070.02H: CCD00M020.00U: FGD002001.00X: > FITLD1: opserver KALI CCD005049.00R; AND001001.05K: CCD00A070.0RQ; CCD00M021.00U: FGD002001.0P0: > FITLD1: Observ. IKALI. AND001001.0II: CCD00504A.00R: CCD00A071.02H: CCD00M022.00U: FGD002001.0RN: > FITLD1: # visibi KALI AND001001.0M8: CCD00504B.00R: CCD00A071.0R0: CCD00M023.00U: FGD002001.1CN: IKALI. > FITLD1: Rand axe AND001001.00W: CCD00504C.00R: CCD00A072.02H: CCD00M024.00U: FGD002001.556: KALI > FITLD1: CCD00A072.0RQ; AND001001.00X: CCD00504D.00R: CCD00M025.00U: FGD002002.00U: KALI > FITLD1: Type AND001001 0P0+ CCD00504F 00R+ CCD000073 02H+ CCD00M026.00U; FGD003001.01D; KALI > FITLD1: COMPLEX ANDOO1C Actual data location - usually no 7.00U; FGD003001.02H: > FITLD1: STOKES KALI ANTIO010 8.00U; FGD003001.00W: ANDOO1C need to look there > FITLD1: FREQ IKALI. FGD003001.00X: 9.000: KALI. > FITLD1: IF 'ANDOO10 A.00U: FGD003001.0P0: KALI > FITLD1: RA ANDO01001.0RS: CCD00504J.00R: CCD00A075.0R0: CCD00M02B.00U: FGD003001.0R0: KALI. > FITLD1: DEC AND001001.0UK: CCD00504K.00R: CCD00A076.02H: CCD00M02C.00U; FGD003001.0UK: KALI > FITLD1: ANDO01001.0XC: CCD00504L.00R: CCD00A076.0R0: CCD00M02D.00U: FGD003001.1CN: KALI > FITLD1: Coordina AND001001.118: CCD00504M.00R: CCD00A077.02H; CCD00M02E.00U; FGD003002.02H: KALI > FITLD1: Rest fre AND001001.1CN: CCD00504N.00R: CCD00A077.0RQ; CCD00M02F.00U; FGD003003.02H: > FITLD1: Alt ref. KALI CCD005040.00R; CCD00A078.02H; FGD004001.05K; AND001001.556: CCD00M02G.00U: KALI > FITLD1: Maximum FGD004001.0G0: AND001002.0RJ: CCD00504P.00R: CCD00A078.0R0: CCD00M02H.00U: KALI > FITLD1: Maximum > FITLD1: Maximum version number of extension files of type BL is KALI 1 > FITLD1: Maximum version number of excapsion files of type FG is KALI 1 > FITL**71**: Appears to have ended successfully KALI > FITLD_t_kali 31DEC09 TST: Cpu= | 0.1 **Seal**= IN= KALI Û. 4

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EUROPEAN ARC	HIPS 1	r;go tkpl∏⊃	scale, non V => pixel scale.



Combining Arrays

- MERLIN+EVN Markarian 273
 - Data already calibrated separately in 2 SPLIT files
 - Check properties in common (check uv data)
 - Phase-ref and pointing position \checkmark
 - Frequency and spectral configuration \checkmark
 - Flux density scale 🗶
 - Correct flux density
 - Select suitable weights, combine data and image
 - Tweak weightings to improve resolution or sensitivity
- AIPS needed to provide user-friendly tasks
 - For rescaling amplitudes & changing weights
 - (essential for delay & rate cal, see Mark's session)
 - Can go back to CASA for advanced imaging





MRK 273 MERLIN+EVN

Calibrate/combine in AIPS, image in CASA
 more options to bring up extended emission



See faint SW emission