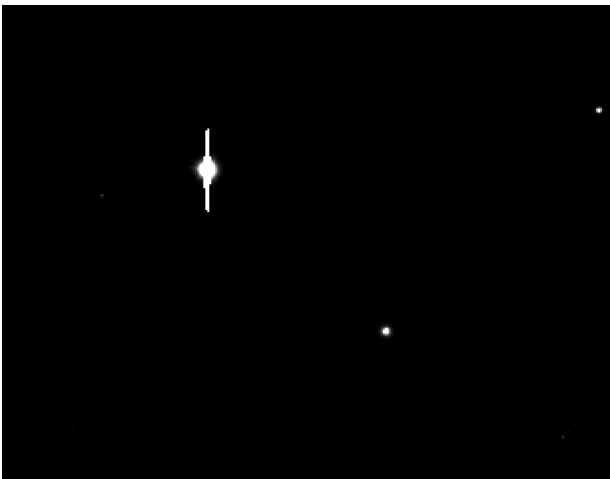
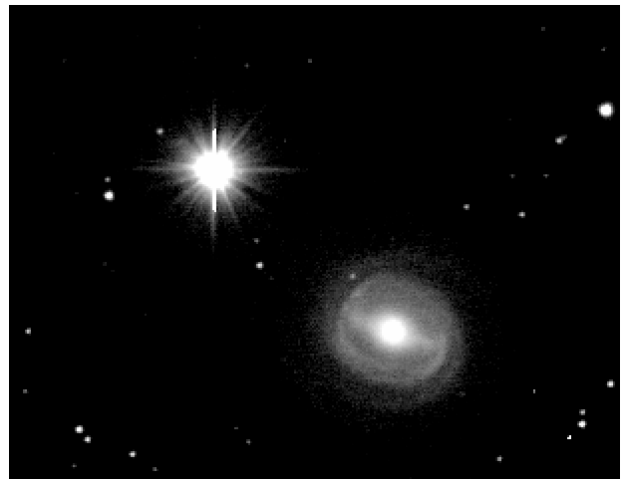


Broad-Band Spectra

Active Galaxies



NGC 3783: *linear* intensity scale

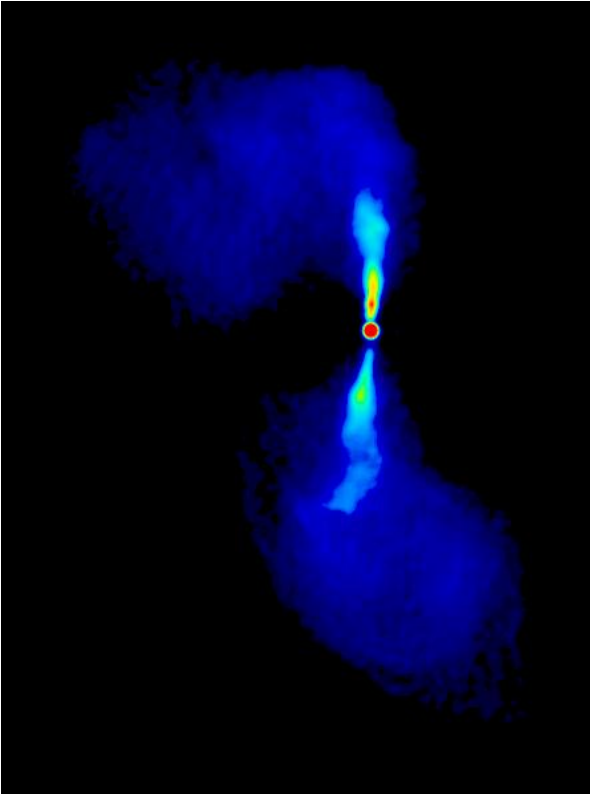


logarithmic intensity scale

Active Galactic Nuclei (AGN): **supermassive black holes** ($M \sim 10^{6...8} M_{\odot}$),
accreting $1 \dots 2 M_{\odot}/\text{year}$

\Rightarrow **Luminosity** $\sim 10^{10} L_{\odot}$ (comparable to galaxy luminosity)

Active Galaxies



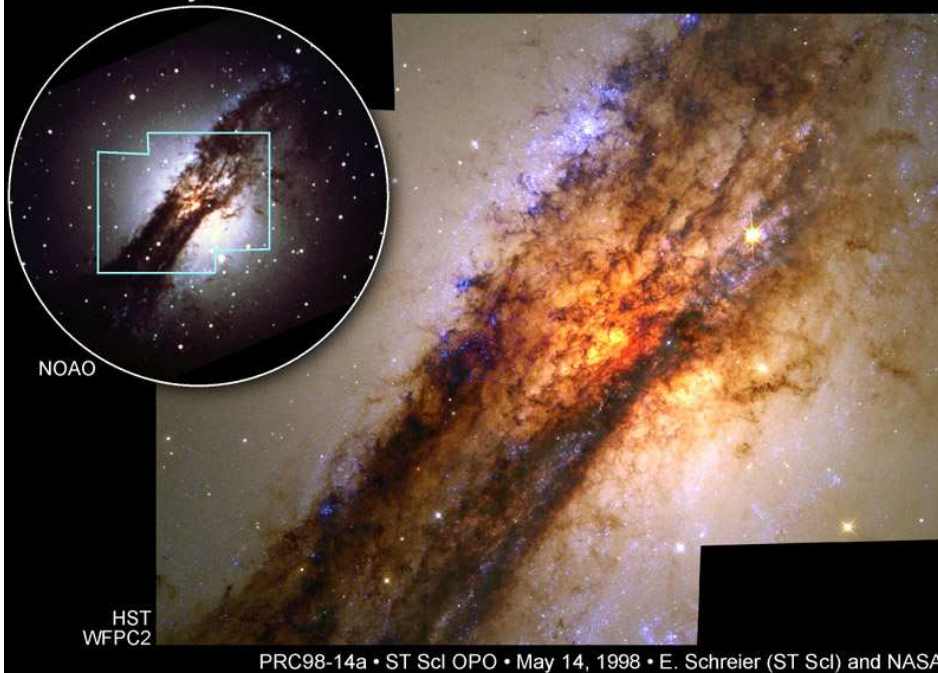
Structure of **Active Galactic Nuclei** (AGN):

- **supermassive black hole** ($10^7 M_{\odot}$)
- **accretion disk** ($\dot{M} \sim 1 \dots 2 M_{\odot} \text{ yr}^{-1}$)
- **large luminosity** ($L \sim 10^{10} L_{\odot}$)
- **Schwarzschild radius** $2GM/c^2 \sim 1 \text{ AU}$
- often **relativistic jets**, where material is accelerated to the speed of light

AGN *with* jets: quasars, blazars. . .

AGN *without* jets: Seyfert galaxies

Active Galaxy Centaurus A



In the following as an example:

Centaurus A (NGC 5128)

- one of the brightest radio sources in the sky
- distance: 11 million light years
- giant elliptical galaxy (more properly: S0), merged with spiral galaxy about 100 million years ago, remnant of the spiral seen as dust lane.

AGN are exceptionally good examples for the importance of multi-wavelength astronomy.

Centaurus A



Optical:

Thermal emission from **stars** and **gas**, i.e., **bremsstrahlung** (free-free radiation), **line emission**, **dust scattering**,...

Cen A: VLT Kueyen+FORIS2, courtesy ESO

Centaurus A



Near Infrared:

Thermal emission, mainly from **stars**, similar to optical, but dust less apparent

⇒ *Opacity* of dust in IR is smaller.

2MASS, courtesy IPAC, Univ. Massachusetts

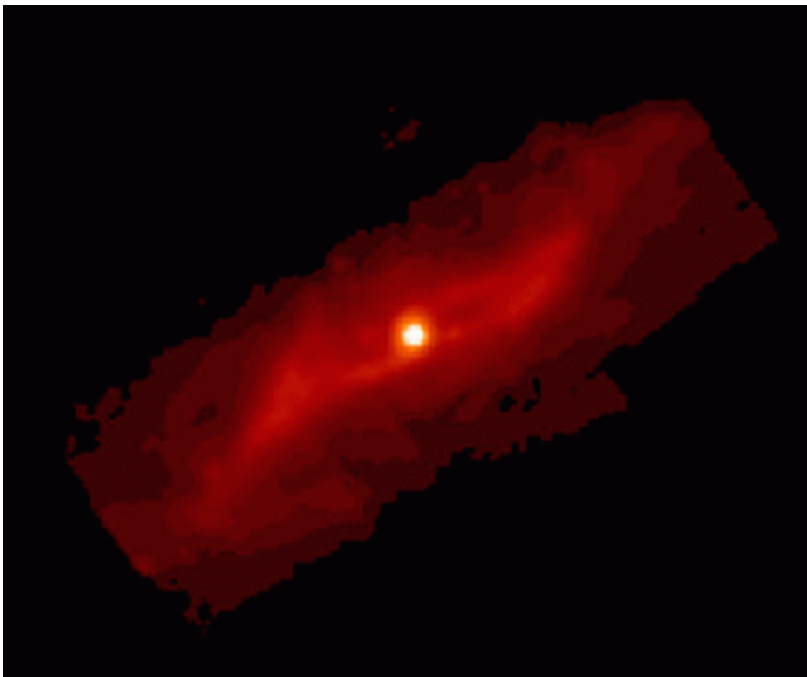
Centaurus A



Mid Infrared
(3.6–8 μm):
Thermal emission
from dust starts to
dominate,
contribution of
thermal emission
from stars still
significant.

Spitzer Space Telescope, courtesy Caltech/NASA

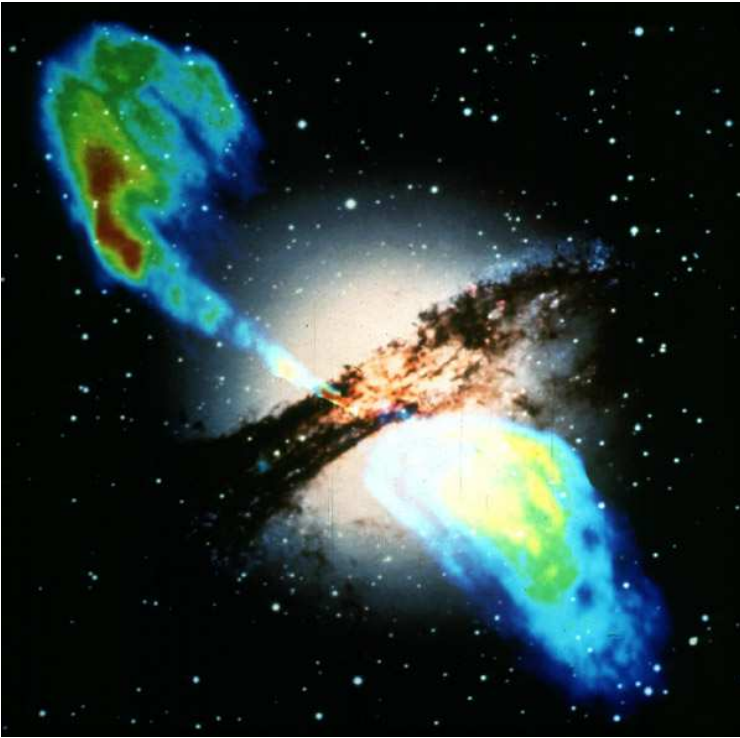
Centaurus A



Far Infrared (7 μm):
Thermal emission from dust
Resolution of this image is worse than
the previous Spitzer telescope image.

ISO, courtesy ESA-ESTEC

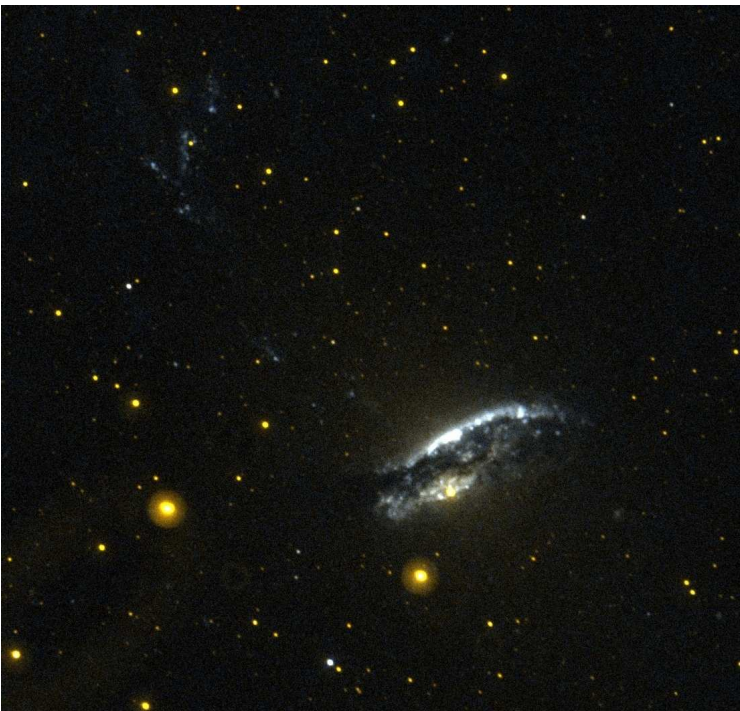
Centaurus A



Radio (6 cm):
 Synchrotron radiation from jets
 and black hole.

VLA/optical, courtesy STScI

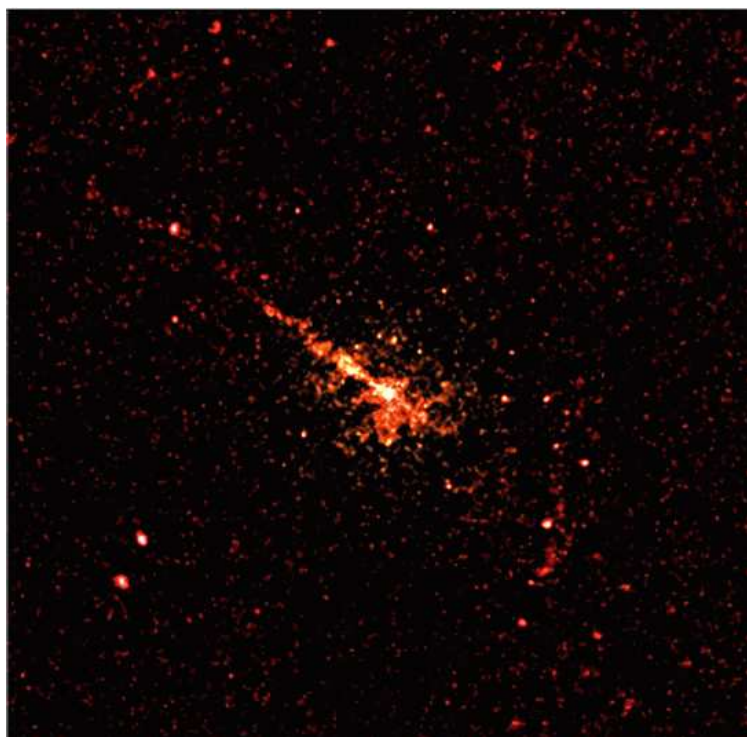
Centaurus A



UV (30–300 nm):
 Thermal UV emission from young
 stars (in NE corner)
 Photoabsorption and absorption
 by dust by dust lane

GALEX, courtesy NASA/Caltech

Centaurus A



Chandra, courtesy CXC

X-rays (2–10 keV):

- Synchrotron radiation from jet,
- Comptonized photons from black hole,
- other emission from X-ray binaries and background AGN

Centaurus A



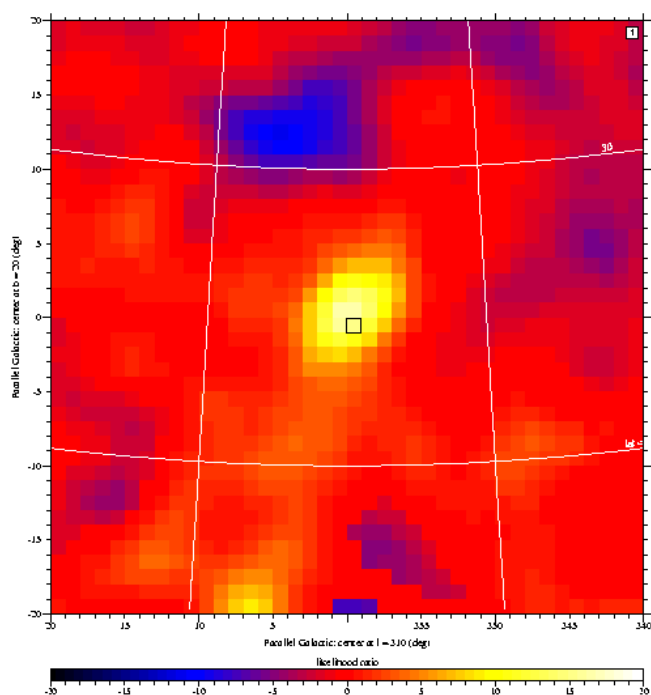
Chandra, courtesy CXC

X-rays (2–10 keV):

- Synchrotron radiation from jet,
- Comptonized photons from black hole,
- other emission from X-ray binaries and background AGN

Centaurus A

Cen A Region: All Phase I+II+III+IV/Cycle 4; 1 - 30 MeV



γ -rays (1–30 MeV):
Comptonized synchrotron radiation
from jet and/or black hole.

CGRO-COMPTEL, courtesy MPE/H. Steinle